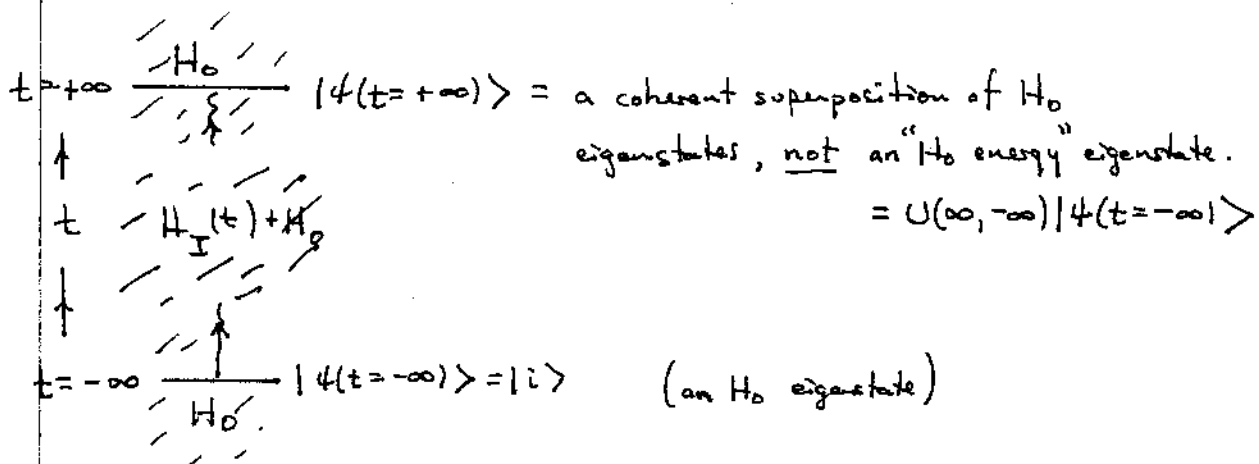


## Time dependent perturbation theory

Often we have a system in which there is a clear division of the Hamiltonian into a fixed  $H_0$  and an interaction  $H_I(t)$  with explicit time dependence, e.g. switched fields or currents,

$$H = H_0 + H_I(t)$$

The quantum problem is then to find the amplitude for a given initial state  $|i\rangle$  at  $t = -\infty$  to evolve into some final  $H_0$  eigenstate  $|f\rangle$ ,  
an eigenstate of  $H_0$



n.b. this is the  $S$ -matrix problem:

$$\underline{S_{fi}} = \underline{\langle f | \psi(t = +\infty) \rangle} = \underline{\langle f | U(\infty, -\infty) | i \rangle}$$

to find  $|\psi(t)\rangle$  with this full  $H$  we solve the TDSE

$$i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle = [H_0 + H_I(t)] |\psi(t)\rangle$$

or with explicit coords

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{x}, t) = [H_0 + H_I(t)] \psi(\vec{x}, t)$$

To develop pert. th. it's useful to expand  $|\psi(t)\rangle$  in a complete set of  $H_0$  eigenvectors,

$$|\psi(t)\rangle = \sum_n a_n(t) |\psi_n(t)\rangle$$

$$\psi(\vec{x}, t) = \sum_n a_n(t) \cdot \psi_n(\vec{x}) e^{-i\omega_n t} \quad \left\{ \begin{array}{l} \\ \rightarrow E_n/\hbar \end{array} \right.$$

substituting this expansion into the Schrödinger equation

$$i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle = [H_0 + H_I(t)] |\psi(t)\rangle$$

& taking the  $\langle m | \cdot \{ \quad \quad \quad \} \text{ overlap}$

gives

$$i\hbar \dot{a}_m(t) = \sum_n \langle m | H_I(t) | n \rangle e^{i(\omega_m - \omega_n)t} a_n(t)$$

Thus far this is completely general. Now suppose that at  $t = -\infty$  we start in the initial state  $|i\rangle$ , so  $a_m(t = -\infty) = \delta_{mi}$ . Also suppose that  $H_I$  is weak, so we can carry out an expansion in  $H_I$ .

The  $\mathcal{O}(H_I)$  term is, setting  $a_m(t) = \underbrace{\delta_{mi}}_{a_m^{(0)}(t)} + \underbrace{\mathcal{O}(H_I)}_{a_m^{(1)}(t)}$ ,

we find

$$i\hbar \dot{a}_m^{(1)}(t) = \langle m | H_I(t) | i \rangle e^{i(\omega_m - \omega_i)t}$$

& integrating,

$$a_m(t) = \underbrace{\delta_{mi}}_{a_m^{(0)}(t)} - \underbrace{\frac{i}{\hbar} \int_{-\infty}^t \langle m | H_I(t') | i \rangle e^{i(\omega_m - \omega_i)t'} dt'}_{a_m^{(1)}(t)} + \mathcal{O}(H_I^2)$$

to calculate the transition probability we now take  $t \rightarrow +\infty$

$$a_i(\infty) = 1 - \frac{i}{\hbar} \int_{-\infty}^{\infty} \langle i | H_I(t') | i \rangle dt' + \mathcal{O}(H_I^2)$$

↑  
still in  $i$

$$P_i(\infty) = |a_i|^2 = \left| 1 - \frac{i}{\hbar} \langle \dots \rangle \right| \left| 1 + \frac{i}{\hbar} \langle \dots \rangle \right| + \mathcal{O}(H_I^2)$$

$$= 1 + \underbrace{\mathcal{O}(H_I^2)}$$

would have to do next order part. try to get this.

amplitude for transition  $i \rightarrow m \neq i$  is

$$a_m(+\infty) = -\frac{i}{\hbar} \int_{-\infty}^{\infty} \langle m | H_I(t) | i \rangle e^{i(\omega_m - \omega_i)t} dt + \mathcal{O}(H_I^2)$$


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$$P_m(+\infty) = |a_m|^2$$

Note: <sup>ΔE</sup> no transitions without time dependence in  $H_I$ ! (In this picture.)

If  $H_I$  is ~ static,

$$a_m(+\infty) = -\frac{i}{\hbar} \langle m | H_I | i \rangle \underbrace{\int_{-\infty}^{\infty} e^{i(\omega_m - \omega_i)t} dt}_{(2\pi) \delta(\omega_m - \omega_i)}$$

Generally there will be time dependence, e.g. external sources or emission/absorption of photons.

In the latter case

$$H_I = - \int d^3x \vec{j} \cdot \vec{A}(\vec{x}, t)$$

leads to  
a matrix element  
of  $\vec{V} \Rightarrow \vec{x}$

plane wave for an EM wave,

$$\vec{A} \propto \hat{e}_{k,\lambda} e^{-i\omega_k t + i\vec{k} \cdot \vec{x}}$$

this will drive  
transitions with  
a  $\Delta E$  of  $\pm \hbar \omega_k$

e.g.



$$\int e^{i(\frac{E_f}{\hbar} - \frac{E_i}{\hbar})t} e^{-i\omega_k t} dt$$

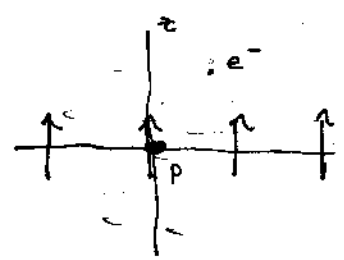
$$\propto \delta(E_f - E_i - \hbar \omega_k)$$

photons of freq.  $\omega_k$   
drive  $\Delta E = \hbar \omega_k$   
transitions.

Thus spectra (F.T.)  
of time dependence is  
crucial for determining  
transition amplitudes.

A simple example of explicit  
time dependence in  $H_I(t)$ .

Suppose we put a bound  $e^-$   
(e.g. in a H atom) in an  
external, modulated electric  
field. What are the transition  
probabilities to the various  
atomic excited states?



$$\vec{E} = E \hat{z} f(t)$$

$$H_I(t) = -eEz \cdot f(t)$$



$$a_{1S \rightarrow 2P} = -\frac{i}{\hbar} \cdot \langle 2P | -eEz | 1S \rangle \cdot \frac{2}{\omega} \sin(\omega T/2) \cdot e^{i\omega T/2}$$

irrelevant phase, depends on when we started the clock.

$F(\omega) =$

$$P_{1S \rightarrow 2P} = \left( \frac{eE}{\hbar} \right)^2 |\langle 2P | z | 1S \rangle|^2 \cdot \frac{4}{\omega^2} \sin^2(\omega T/2)$$

to be effective in exciting transitions, the time scale of our pulse should be

$$T \sim \frac{\pi \hbar}{\Delta E}$$

our case

$$\Delta E = E_{2P} - E_{1S} = \frac{1}{2} mc^2 \alpha^2 \left( 1 - \frac{1}{4} \right) = \frac{3}{8} mc^2 \alpha^2 = 10.2 \text{ [eV]} \quad \text{(soft X-rays!)}$$

$$\therefore \text{need } T \sim \frac{\pi \hbar}{\Delta E} = \frac{3\pi}{8} \frac{\hbar}{mc^2 \alpha^2} = \frac{3\pi}{8} \frac{1.054 \cdot 10^{-27} \text{ [erg} \cdot \text{sec]}}{511.0 \text{ [keV]}} \cdot (137.036)^2$$

$$\sim 1 \cdot \frac{10^{-27} \text{ [erg} \cdot \text{sec]} \cdot 2 \cdot 10^4}{\frac{1}{2} \cdot 10^6 \cdot 1.6 \cdot 10^{-12} \text{ [erg]}} \sim \frac{2 \cdot 10^{-23} \text{ [sec]}}{10^{-6}}$$

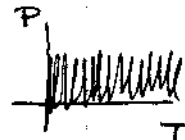
$$T \sim 2 \cdot 10^{-17} \text{ [sec]}$$

So, our switch is much too slow to drive this transition effectively.

for our case  $\langle 2P | z | 1S \rangle = \frac{2^{15/2}}{3^5} a_0$ ,

$$P_{1S \rightarrow 2P} = \frac{2^{17}}{3^{10}} \left( \frac{eEa_0}{\hbar} \right)^2 \underbrace{\frac{1}{\omega^2} \sin^2\left(\frac{\omega T}{2}\right)}$$

rapidly osc. factor since  $T \gg \omega^{-1}$   
replace by  $1/2$  ( $\omega T \sim 10^{11}$ )



$$\overline{P}_{1S \rightarrow 2P} = \frac{2^{16}}{3^{10}} \left( \frac{eEa_0}{\hbar\omega} \right)^2$$

$$\hbar\omega = E_{2P} - E_{1S} = 10.2 \text{ [eV]}$$

$$eEa_0 = 10^4 \text{ [eV cm}^{-1}] \cdot 5.293 \cdot 10^{-9} \text{ [cm]} \\ = 5.29 \cdot 10^{-5} \text{ [eV]}$$

$$\therefore \left( \frac{eEa_0}{\hbar\omega} \right)^2 = \left( \frac{5.29 \cdot 10^{-5} \text{ [eV]}}{10.2 \text{ [eV]}} \right)^2 = 2.69 \cdot 10^{-11}$$

$$\overline{P}_{1S \rightarrow 2P} = 2.99 \cdot 10^{-11}$$

$\therefore$  A few of the H atoms are excited! (3 in a trillion, but even so this is  $3 \cdot 10^{-11} \cdot 6 \cdot 10^{23} \sim 2 \cdot 10^{13}$  per mole)

n.b. even so, this is unphysical!  $10^{13} \text{ } \overset{10\text{eV}}{\text{X-rays}}$  emitted by the sample when the 2P states decay in the next  $\mu\text{sec}$ !

This fiction is caused by the hard edge of our assumed pulse which gives  $a \sim \frac{1}{\omega}$ ,  $P \sim \frac{1}{\omega^2}$  dependence in the ultraviolet.



The real pulse will be much smoother at the  $\frac{\hbar}{\Delta E} \sim 10^{-17}$  [sec] time scale, leading to a much smaller probability of excitation.

No square pulses in nature!

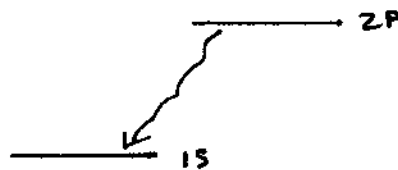
Fermi's Golden Rule (d radiative transitions)

Thus far we have considered time dependence of transitions between discrete levels, which we can get by just solving the matrix equation

$$i\hbar \frac{\partial}{\partial t} |\psi\rangle = H |\psi\rangle$$

next we consider the transition rate from an initial state  $|i\rangle$  into a continuum of states  $|f\rangle$ .

(e.g. radiative transitions,



$|i\rangle = |2P\rangle$  discrete level

$|f\rangle = |1S\rangle | \gamma_{k,\lambda}^+ \rangle$

↳ there is a continuum of single photon states.

for discrete levels, (Born-order)

$$a_{f \neq i}(\infty) = -\frac{i}{\hbar} \int_{-\infty}^{\infty} \langle f | H_I(t) | i \rangle e^{i(\omega_f - \omega_i)t} dt$$

has a  $\langle f | H_I | i \rangle e^{i\omega_f t}$  part

$$= -\frac{i}{\hbar} \langle f | H_I | i \rangle 2\pi \delta(\underbrace{\omega_{1S} + \omega_{\gamma}}_{\substack{\uparrow \\ \omega_f, \\ \text{part of a} \\ \text{continuum}}} - \underbrace{\omega_{2P}}_{\substack{\uparrow \\ \omega_i, \\ \text{discrete}}})$$

$$P_{f \neq i}(\infty) = \frac{1}{\hbar^2} |\langle f | H_I | i \rangle|^2 2\pi \delta(\omega_{1S} + \omega_{\gamma} - \omega_{2P}) \cdot T$$

$$\delta(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{i\omega t} dt$$

$$\Gamma_{i \rightarrow f} = \frac{P}{T} = \frac{1}{\hbar^2} |\langle f | H_I | i \rangle|^2 \cdot 2\pi \delta(\omega_{1S} + \omega_{\gamma} - \omega_{2P})$$

$$\delta(\omega)^2 = \delta(\omega) \delta(0)$$

$$\delta(0) = \frac{T}{2\pi}$$

$$\therefore [2\pi \delta(\omega)]^2 = 2\pi \delta(\omega) \cdot T$$

The transition rate to a set  $\{f\}$  of distinguishable final states is

thus no interference

$$\Gamma_i = \sum_f \Gamma_{i \rightarrow f}$$

which we can replace by an integral over the final continuum,

$$\sum_f = \sum_f \underbrace{\Delta n}_{\substack{\text{counting through} \\ \text{a list } n=1 \dots N \text{ of final states}}} = \sum \frac{\Delta n}{\Delta \omega} \Delta \omega \xrightarrow{\text{continuum limit}} \int \underbrace{\frac{dn}{d\omega}}_{\substack{\text{density of final states,} \\ \text{dn between } \omega \text{ and } \omega + d\omega, \\ \rho(\omega) = \frac{dn(\omega)}{d\omega}}} d\omega$$

So, for transitions into a continuum,

$$\sum_f \Gamma_{i \rightarrow f} = \frac{1}{\hbar^2} |\langle f | H_I | i \rangle|^2 2\pi \int \delta(\omega_f - \omega_i) \rho(\omega_f) d\omega_f$$

n.b.  
 $d\Gamma_f = \frac{2\pi}{\hbar} |\langle f | H_I | i \rangle|^2$   
 $\cdot \rho(E_f) \frac{dE_f}{4\pi}$   
 into  $d\Omega$  if m.e. is not isotropic

$$\Gamma_{i \rightarrow f} \text{ (in continuum)} = \frac{1}{\hbar^2} |\langle f | H_I | i \rangle|^2 \cdot 2\pi \rho(\omega_f)$$

Fermi's Golden Rule

$\frac{2\pi}{\hbar} |\langle f | H_I | i \rangle|^2 \rho(E_f)$

units?  $[\text{sec}^{-1}] = \frac{1}{[\text{erg}^2 \text{sec}^2]} \cdot [\text{erg}^2] \cdot \left[ \frac{1}{\text{sec}^{-1}} \right]$

$= [\text{erg} \text{sec}] \cdot [\text{erg}^2] [\text{erg}^{-1}]$

Actually  $d\Gamma_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle f | H_I | i \rangle|^2 dn$